



# Gammapy Python package for gamma-ray astronomy into the Open Science

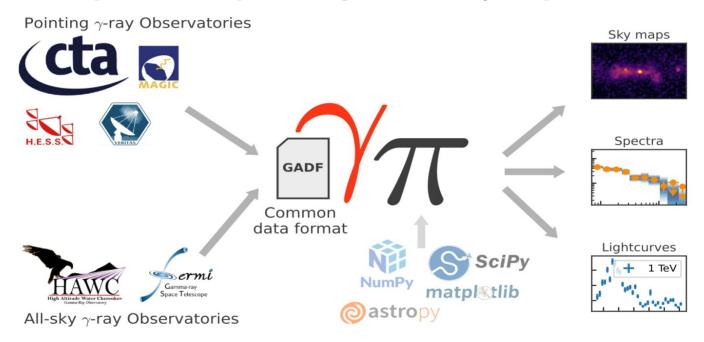
Bruno Khélifi Project manager, for the Gammapy team

RICAP-24, Villa Tuscolana, Frascati September 26<sup>th</sup>, 2024



## **Open Science Tool: its concept**

## Python library to analyse high-level $\gamma$ -ray data



Designed to analyse several data sets

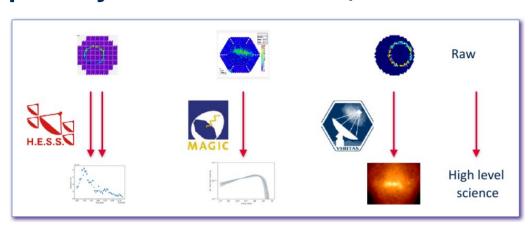






## The landscape 10 years ago

#### Proprietary data and formats, closed software tools



VHE community worked in a totally competitive and closed mode

 Except few MoUs around scientific projects

# Some 'dreamers' worked towards the opening of the VHE astrophysics

- Data format standardization: open initiative
   'Open Gamma-Ray Astro' → GADF format
- Open Science Tools: <u>Gammapy</u>

Deil, C., et al., ASTERICS 2016 (link)

Better results
Interoperability between instruments
Respect of the FAIR principles







## Early steps of Gammapy: 2014 – 2017

- Github repository creation in August 19th, 2013 (TevPy)
- First Gammapy release (v0.1) on August 25th, 2014
- Project evolved into a generic library for TeV astronomy and in prevision of the CTAO science analysis tool

« We would like to introduce Gammapy to the community and present our vision of Gammapy as a future community-developed, general purpose analysis toolbox for  $\gamma$ -ray astronomers. [...] Its scope will continuously grow and we hope that many users and developers show interest in open and reproducible  $\gamma$ -ray astronomy with Python. As long-term goal we would like Gammapy to turn into a fully community-developed package. »



Donath, A., Deil, C. et al. ICRC 2015



## **Towards the LTS v1.0 : 2018 - 2022**

- Rapid development cycle with frequent releases (~ 2 month)
  - From v0.7 to 0.20
- Structuration of the library & abstraction of analysis steps
- 19,000 commits from more than 80 contributors
- June 2021: Gammapy selected as official CTAO Science Analysis Tool
  - Used for the CTAO Real Time Analysis
- Version 1.0 released Nov. 10<sup>th</sup>, 2022

See v1.0 Gammapy paper: Donath et al. (2023)

#### PIG 5 - Gammapy 1.0 roadmap

- · Author: Axel Donath, Régis Terrier & Christoph Deil
- Created: Sep 28, 2018
- Accepted: Jan 31, 2019
- · Status: accepted
- Discussion: GH 1841

#### **Abstract**

This PIG describes the required short- and medium-term **development work up to the Gammapy 1.0** release. The anticipated time scale for this development effort is 9 - 12 months
and will be concluded by the Gammapy 1.0 release in fall 2019. The question of API design
and sub-module structure for Gammapy 1.0 will be addressed in separate PIGs.

The content of this document was decided based upon user feedback from the first CTA data challenge (DC1), experience from analysing existing datasets as well as definition of use cases (see below). The content will be updated in the coming month and be adjusted to upcoming requirements defined by CTA. Current requirements defined by CTA are described observer access use cases (private link to sides) and in the document written summarizing the SUSS workshop Dec. 2018 (private link to indico).









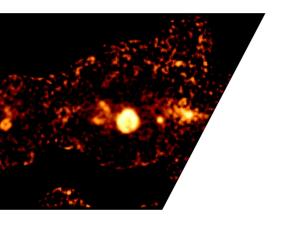
### From 2014 to 2024



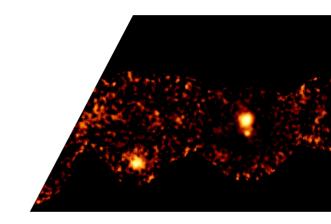
#### Talk content:

- 1- Design and features
- 2- Joint multi-instrument analyses
- 3- An open science project





## **Design and features**





## The Gammapy library

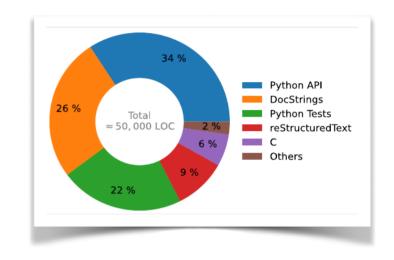
#### **Lightweight Python software**

- Astropy-affiliated package
- Fermipy depends on Gammapy

#### **Openly developed on Github**

- 8-10 core contributors
- More than 80 contributors from the whole  $\gamma$ -ray astronomy community and beyond

### Distributed via PyPi and conda-forge













## Data analysis workflow & package structure

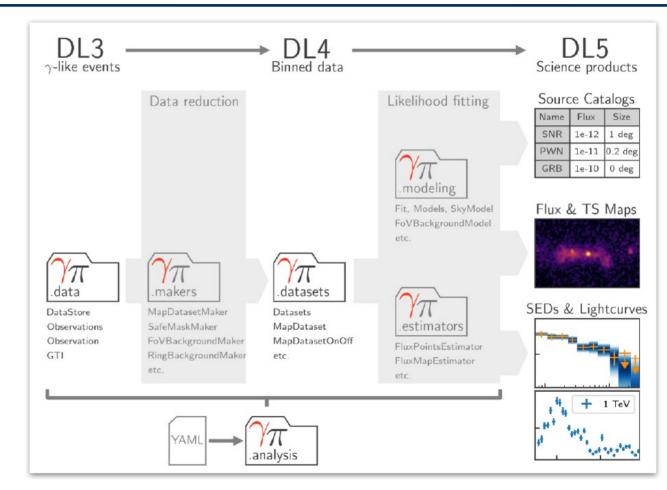
#### 2-step analysis procedure:

- data aggregation and reduction (DL3 to 4)
- modeling / fitting (DL4 to 5)

Allow for joint data modeling at DL4 level

#### Flexible modeling library:

- physical models (e.g. naima)
- user designed models

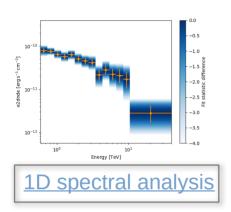


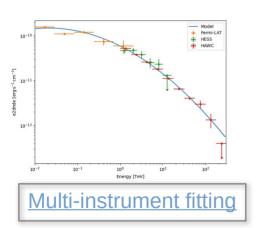


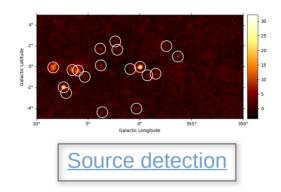


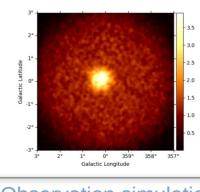


## Typical analysis use cases

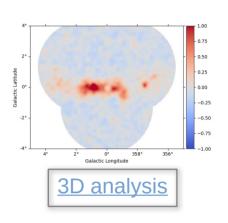


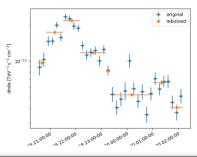












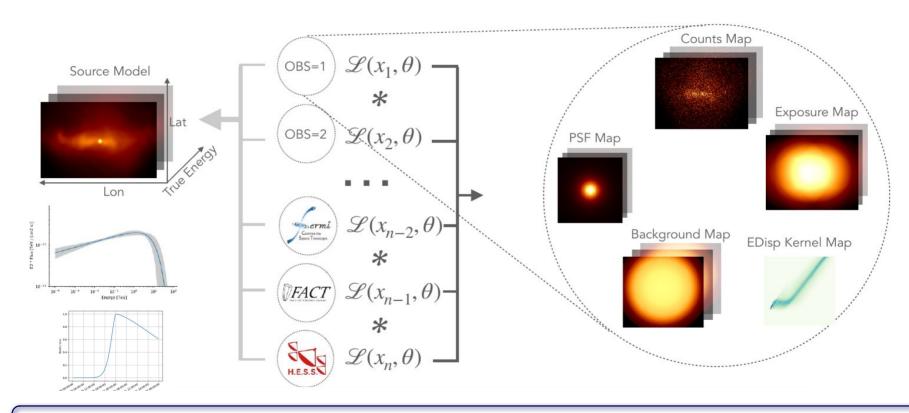
<u>Light-curve extraction</u> and <u>time-variability estimation</u>







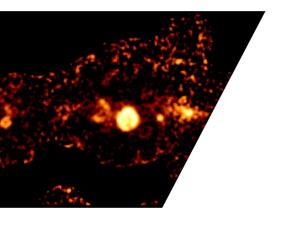
## Joint likelihood and multi-instrument analysis



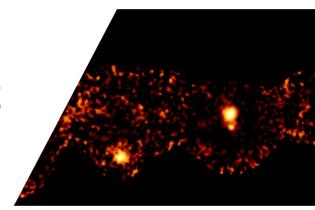
Gammapy Dataset structure allows heterogeneous data fitting. See joint fit tutorial







# Joint multi-instrument analyses



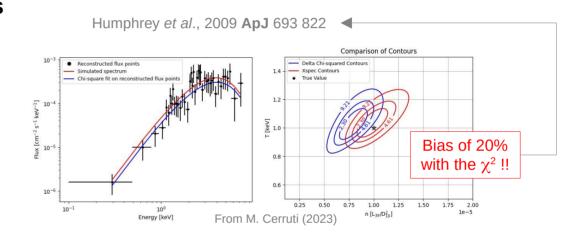


## Challenges for multi-instrument analyses

#### Rigorous data analysis

- Need correct handling of statistics
  - In contrary to basic  $\chi^2$  fit on flux points!

Ex: simulation of a BB seen by Swift with XSPEC



- And inter-instrument systematics
  - E.g. use of priors or parameters or IRFs

#### Readability of the IRFs

- IRFs from HE → UHE instruments can be factorised in the same manner
- Need of the use of standard formats! (GADF → VODF)

Use of 3D analysis



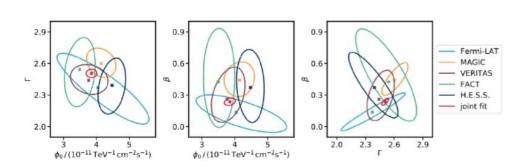




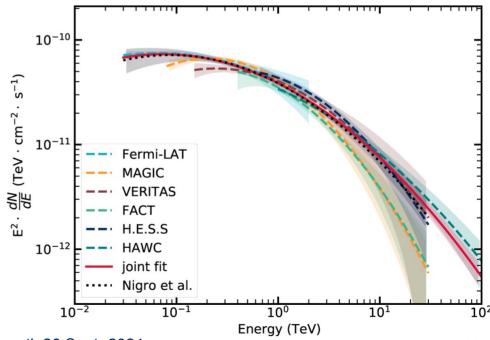
## An example: from HE to UHE $\gamma$ -rays

#### Joint point-like 1D spectral analysis of the Crab nebula

- 6 different instruments over 3.5 decades in energy
  - Simple log-parabola & physical inverse Compton model
  - Modelling of some systematic uncertainties
- Fully reproducible analysis



Nigro et al. 2019 + HAWC: Albert et al. 2022







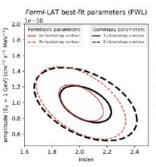


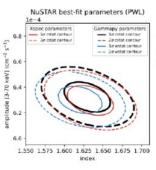
## Other example: from optical to GeV

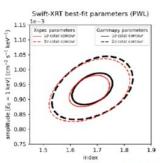


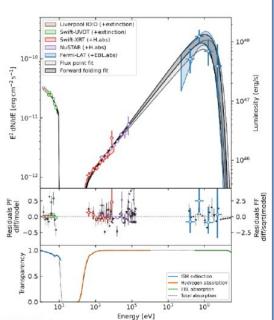
## Multi-instrument analysis over 10 decades

- Joint forward fitting fit from eV to 10<sup>10</sup> eV with Gammapy: Liverpool OT, Swift-UVOT, Swift-XRT, NUSTAR, Fermi-LAT
- Flux points lose some stat. information (e.g UL)
- Full forward fit provides more accurate results
- Gammapy facilitates the distribution and reproducibility of the results









OP 313 campaign

From R. Terrier, Gamma2024









## An other one: from X-rays to VHE $\gamma$ -rays



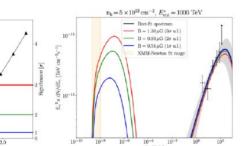
### **Multi-instrument analysis examples**

#### Joint X-ray and γ-ray fits

#### First approach:

Read OGIP spectra (1D DL4)
 produced by X-ray telescopes and fit

gammapyXray package



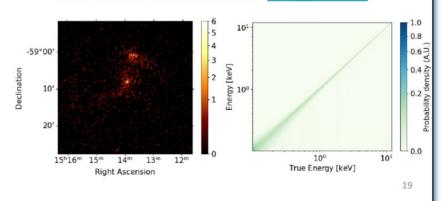
Giunti et al (2022)

DOI 10.5281/zenodo.7092736

#### Second approach:

 Read X-ray events, IRFs and create 3D DL4 dataset

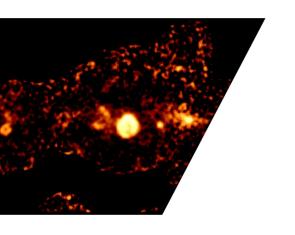
See e-ROSITA converter in K. Egg poster



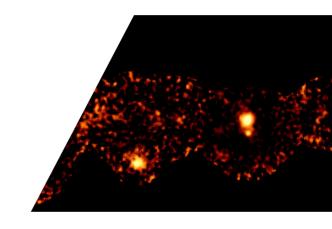
From R. Terrier, Gamma2024







# An Open Science project





## **Open research software**

#### **Recognition and valorization**

- Each release is a real publication
  - According to Open Science recommendations on the science evaluation
  - DOI and SWHID
- Transparent Gammapy Authorship Policy

#### Long-term archive

 On the universal archive Software Heritage (sustained by UNESCO)

## zenodo







#### **Open science activities**

- VHE standards : ← creation of the "High Energy IG" [IVOA, Malta, Nov'24]
- VHE data format : GADF → VODF
- Support of any open project that can be affiliated to Gammapy







### **Dissemination and outreach**

#### **Presentations, hands-on sessions**

Material available here and here

#### **Schools**

- ORP school on Multi-messenger Astrophysics, Durham University, Sept 2-6
- CTAO school, La Palma June 22 29
- MPIK-CDY school on the future of γ-ray astronomy, MPI-K HD, June 25-July 3
- IFSC school, São Marcos Feb./March

Advertisement: Cherenkov Astronomy Data School (CADS)

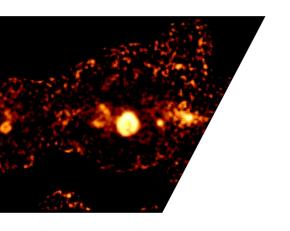
beginner and advanced hands-on sessions

Observatoire de Paris, October 14-18 2024, more on <a href="https://indico.obspm.fr/event/2480/">https://indico.obspm.fr/event/2480/</a>

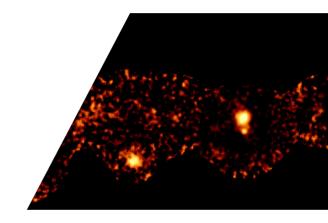
#### Within the collaborations

Dedicated support and training





## **Conclusions**





## Towards the LTS v2.0 and beyond

#### **Towards ML & MM data analysis**

- UHE (HAWC → SWGO)
- X-ray handling improvement
- Neutrino processing
- Unbinned likelihood analysis

#### Better respect of s/w standards

- FAIR4RS principles (towards reproducibility)
- Some IVOA standards (e.g. Provenance)

#### Improved performance

- Computing time (optimisation, JAX?)
- Memory (caching, sparse array, etc)

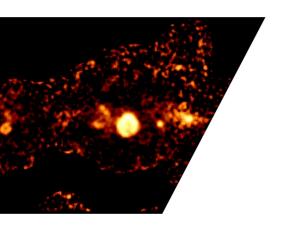
# Get involved in the adventure

2 post-doc positions be announced soon

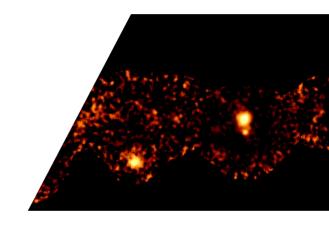








## **Back-up slides**





## **Project Organization**

#### **Projects managers**

non-technical executive lead

#### Lead developers

technical executive leads

#### **Sub-packages maintainers**

core developer devoted to the maintenance of some sub-packages

#### Contributors

>80 individual contributors from various collaborations and beyond

See <a href="https://gammapy.org/team.html">https://gammapy.org/team.html</a>

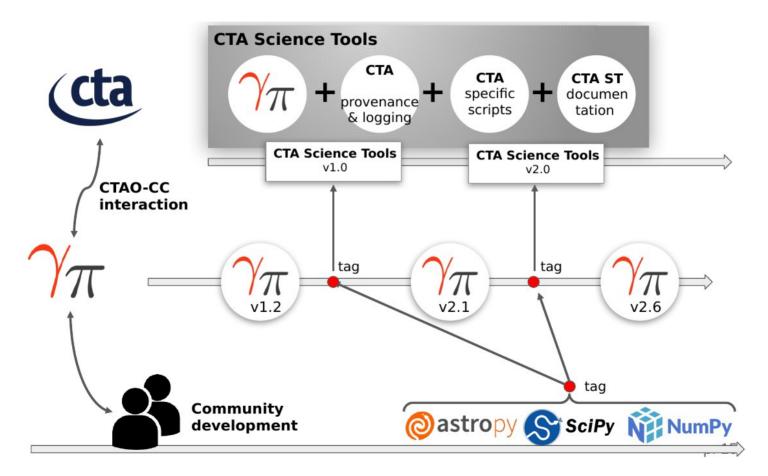
# **Coordination Committee** « Promotes, coordinates and steers Gammapy developments » UNIVERSIDAD COMPLUTENSE MADRID INAF







## **Gammapy and CTAO SAT**







## Towards the LTS v2.0: features selection

- <u>Parameter prior</u> can now be defined on parameters and the associated log-prior is added to the total statistics during fitting
- Added timing studies utility functions for light curves, see <u>tutorial</u>
- Preliminary support for asymmetric IRFs. See tutorial
- Energy dependent ON-region size for 1D spectral analysis
- And many more...
- More to come in v1.3 (in October!) and beyond
  - Improved support for event types and joint analyses
  - Multi-parameters priors and spectral unfolding
  - Lightcurve simulation and PSD study tools
  - ...



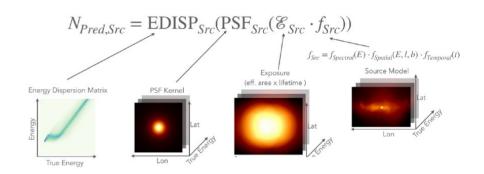


## Poisson Log-Likelihood

## Common algorithms for the libraries: Poisson Log-Likelihood

"Cash statistics": summed over all "bins" 
$$\mathcal{C} = 2\sum_{i} N_{Pred}^{i} - N_{Obs}^{i} \cdot \log N_{Pred}^{i}$$
 i: spectral channels or 3D voxels 
$$N_{Pred} = N_{Bkg} + \sum_{Src} N_{Pred,Src}$$

- Bins in the spectral, spatial, temporal domain
- → Need of a "global" background model template with "correction parameters"



→ Need of the "signal" IRFs and source models

Most of the time,

Identical factorization of the IRFs
for X-rays → UHE & neutrino exp.







## **FAIR4RS** principles

As with the FAIR Guiding Principles, the <u>FAIR4RS Principles</u> (2022) are intended to be aspirational. The application of the FAIR4RS Principles is the responsibility of the owners (who are often the creators) of the software, not the users.

- Software, and its associated metadata, is easy for both humans and machines to find
- Software, and its metadata, is retrievable via standardized protocols
- Software interoperates with other software by exchanging data and/or metadata, and/or through interaction via application programming interfaces (APIs), described through standards.
- Software is both usable (can be executed) and reusable (can be understood, modified, built upon, or incorporated into other software)







#### **VHE data format: VODF**



## Open Initiative 'Very-high-energy Open Data Format' (link)

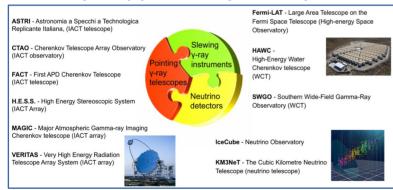
Aims to format VHE data (gamma and

neutrino)



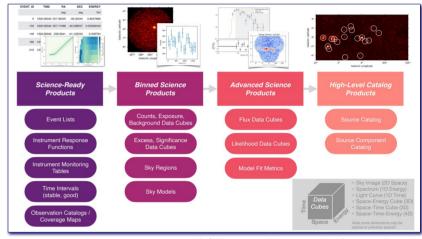


Officially supported by 11 experiments



Structured with a project organization

Coordination Committee, Conveners: R. Zanin, B. Khélifi Lead Editors: K. Kosack, L. Olivera-Nieto, J. Schnabel



Khélifi, B., et al., Proc. of 38th ICRC (2023)







## Getting help / giving feedback - And get involved!!

- Where/How to interact with dev team and experienced users, provide feedback, get help:
  - gammapy.slack
    - In particular: #help channel
  - GitHub discussions
    - · help category
  - GitHub issues to report bugs or feature requests



- Development guide
- The dev calls: each Friday at 14h CEST!
- The <u>hands-on sessions and schools</u>, the <u>recipes</u>
- Etc







## **Getting the software**

Quickstart installation with conda (ex: LTS)

```
curl -0 https://gammapy.org/download/install/gammapy-1.0.2-environment.yml conda env create -f gammapy-1.0.2-environment.yml conda activate gammapy-1.0.2
```

Installation with pip

pip install gammapy



install all dependencies

Download tutorials & associated data

gammapy download notebooks
gammapy download datasets
export GAMMAPY\_DATA=\$PWD/gammapy-datasets



See:

